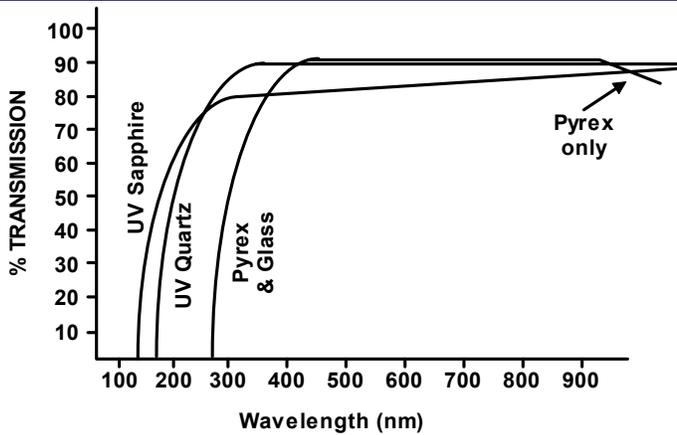


# Applications Information

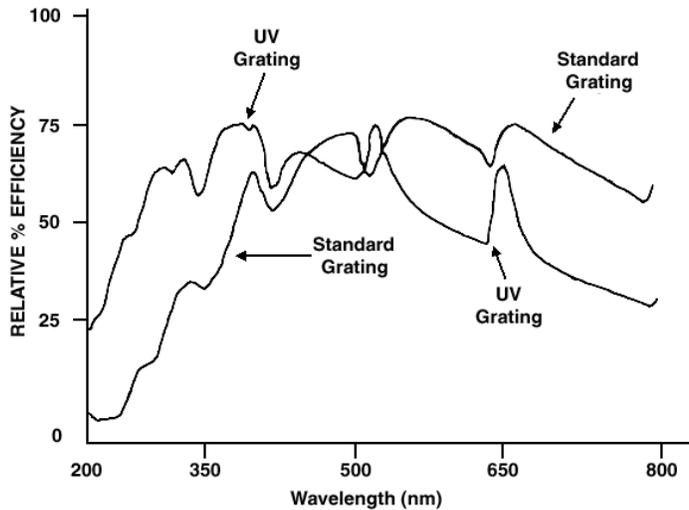
## Window Materials



In applications where plasma is viewed through a window, the proper selection of window material is essential to viewing the desired spectra. Use the chart at the left to select the proper window material, viewports, lenses, etc. As a rule, costs increase as ultraviolet (UV) transmittance improves.

Resistance to etching is an additional consideration in window material selection. Of the options presented here, UV sapphire offers the greatest resistance to etching. Of course, window material selection is not an issue when using a viewing technique that employs a vacuum feedthrough.

## Grating Options



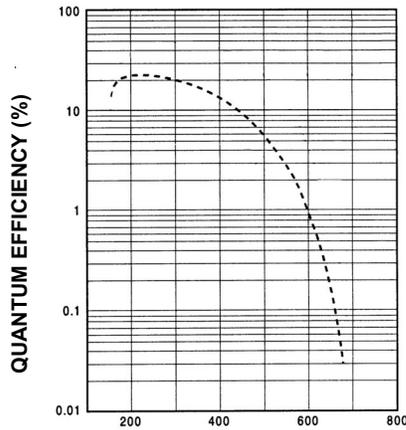
The EP Series monochromators are available with a UV or standard grating. The diagram on the left shows the relative efficiency of these gratings. Note that the blaze wavelength occurs on the low side of the grating  $\geq 30\%$  efficiency range.

Grating Type	Standard	UV
Typical Range	VIS/NIR	UV/VIS
Blaze Wavelength	450nm	350nm
$\geq 30\%$ Efficiency Range	310–920	220–750

# Detector Spectral Response

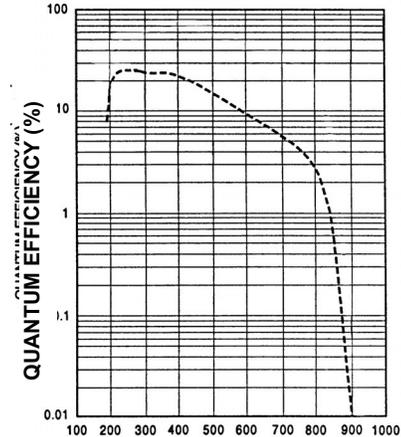
EP200 Series/SD100

### UV Enhanced PMT



WAVELENGTH (nm)  
Range: 185–650nm  
Max. Voltage: 1000VDC

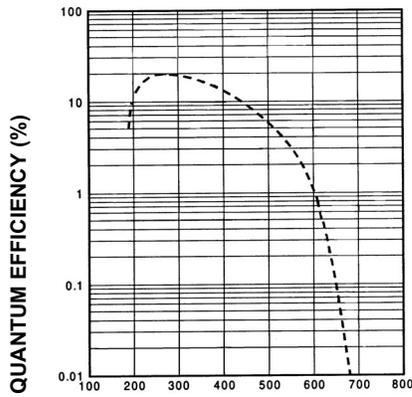
### Standard PMT



WAVELENGTH (nm)  
Range: 185–900nm  
Max. Voltage: 1000VDC

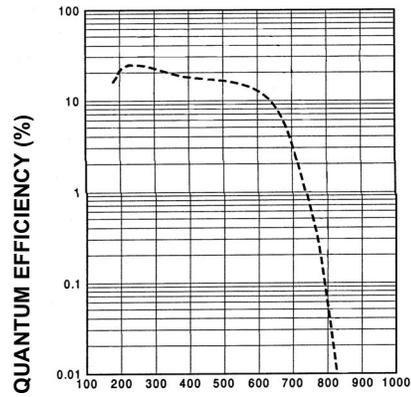
PM Series – PM100/140/220/225

### UV/VIS PMT



WAVELENGTH (nm)  
Range: 185–650nm  
Max. Voltage: 1000VDC

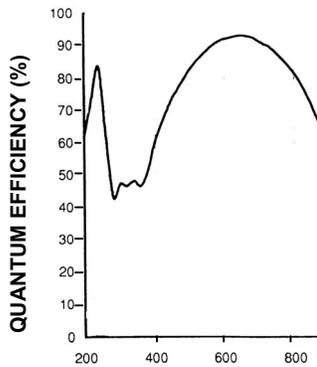
### UV/VIS/NIR PMT



WAVELENGTH (nm)  
Range: 185–830nm  
Max. Voltage: 1000VDC

SD1024D

### 1024 Pixel CCD

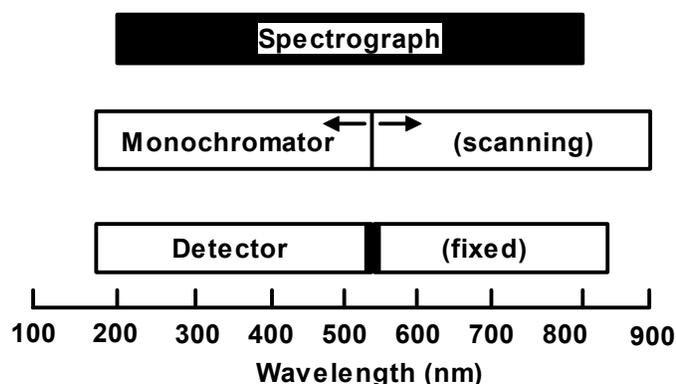


WAVELENGTH (nm)  
Range: 200–800nm

The CCD used in the SD1024D spectrograph provides scientific grade performance. Since it is backthinned, it has excellent UV performance and is TE cooled for increased stability. Additional sensitivity is derived through software that permits the selective addition of discrete wavelengths.

Note: the quantum efficiency of the general-purpose CCD in the SD1024DL and SD2048DL is not represented.

## Instrument Ranges



The wavelength(s) of interest are a major consideration when selecting the proper instrument. Spectrographs, using CCD technology, permit continuous measurement of all wavelengths in the instrument's operating range.

Monochromators measure a single wavelength by means of slits, gratings, mirrors and photomultiplier tubes (PMTs). To change a monochromator's wavelength setting, a stepper motor or micrometer is used to adjust grating position.

Interference filter-based detectors are the simplest, least expensive optical emission analysis instrument. The drawback of the interference filter detector is its wider bandpass, typically 10nm, and the inability to easily change the wavelength analyzed.

### Automatic Gain Control

All Verity photomultiplier-based instruments include a provision for automatic PMT high voltage control. The setting of the PMT's high voltage control determines its sensitivity or gain. When used with Verity's system controllers, or integrally with the SD100, the automatic gain control (AGC) feature implements a gain correction algorithm to automatically adjust PMT high voltage.

The AGC is completed once the output of the PMT detector reaches a preset target value, typically about 7V on a 0–10V scale. When it is found that the signal level is not significantly changing, an automatic gain correction is made at a fixed time interval near the beginning of each process step. In this manner, both the short- and long-term effects of signal degradation can be corrected, providing each endpoint trace starts at the same preset value.

### Correction for Clouding Optics

Clouding of the optical interface can be minimized through the use of a bakeable capillary cartridge assembly as part of a vacuum feedthrough system. See page 79 in the Optics section for additional information on the capillary cartridge assembly.

## Direct Viewing



**Monochromator with Direct Viewing Quartz Lens Coupler**

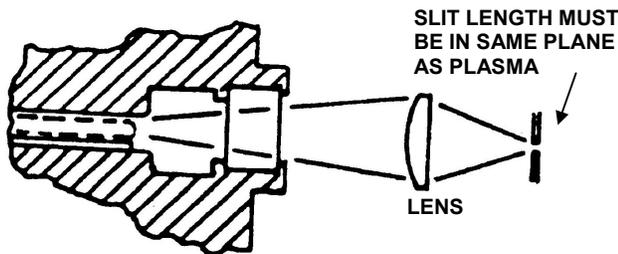
Direct viewing refers to the use of an instrument without a fiber optic cable.

### General Information

Use direct viewing in applications when an optical window provides line-of-sight access to the plasma and when there is room to mount the instrument in close proximity to the window. For this technique, it is essential for the window to possess proper transmission characteristics.

When using the EP200 Series monochromator or the PM100 detector in this viewing mode, please note that harmful UV light could be transmitted through some window types (e.g., quartz and sapphire). A light shield or UV blocker, such as acrylic, should be adapted around the detector bus coupler to prevent UV light leakage outside the detector system.

Direct viewing is standard with PM100, PM140, PM220 and PM225 detectors, and optional with EP200 Series monochromators. Note, however, that the PM140 and PM225 are optically sealed using a KF40 flange, thus eliminating concerns of UV leaks at the instrument/chamber interface.



### Monochromator Slit Orientation

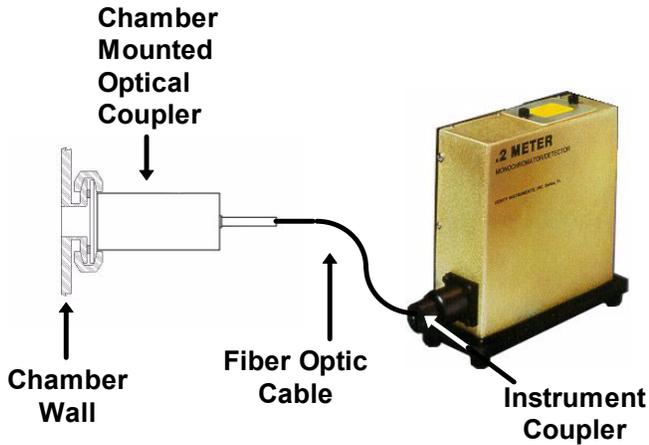
When using direct viewing techniques, it is important that the monochromator slit is oriented in the same plane as the plasma. This orientation will provide the strongest signal for endpoint species. Note that the monochromator's slit length is vertically oriented when the monochromator is placed on the front or back side (the narrow sides, not the largest sides).

## Indirect Viewing

Indirect viewing refers to the use of an optical instrument with a fiber optic cable.

### Without a Vacuum Feedthrough System

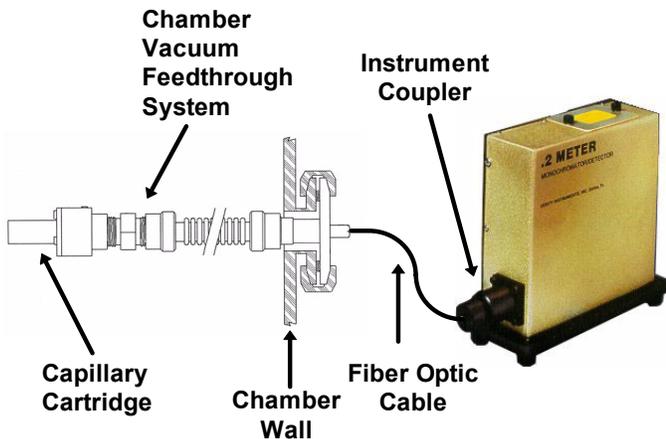
Use indirect viewing without a vacuum feedthrough when an optical window is available with direct viewing access to the plasma, but without room to mount the instrument adjacent to the window. With this technique, it is important that the window have proper transmission characteristics. Additional hardware required for indirect viewing adds to the cost of the system.



### With a Vacuum Feedthrough System

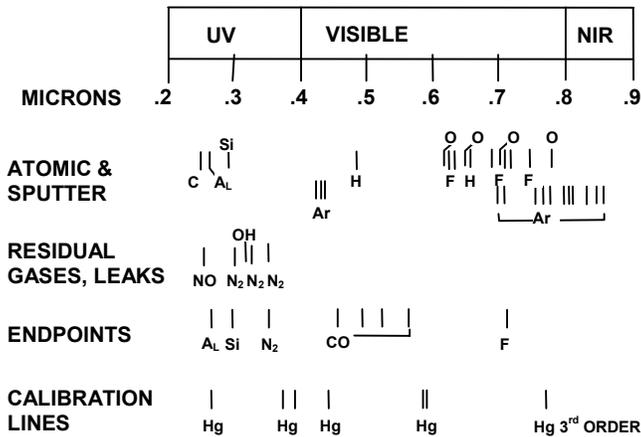
Use indirect viewing with a vacuum feedthrough system in applications where the window or viewport may become coated with process material, or if a feedthrough provides better viewing of the plasma.

This system includes a vacuum feedthrough, flexible tube and an optical head. The optical head includes an optional capillary cartridge, which is an array of small bore stainless steel tubes with a high aspect ratio (length to diameter) that help prevent contamination from reaching the lens. See page 79, Chamber-Mounted Vacuum Feedthroughs for additional information.



# Emission Species

## Spectra of Common Species



## Representative Emission Lines Used in Endpoint Detection of Plasma Etching

Use the table below when considering wavelengths to monitor. This table is an abbreviated list of the wavelength information presented on pages 11 and 12, and has been developed from Verity research and Herman<sup>1</sup>.

Monitored species	Wavelength (nm)
Al	308.2, 309.3, 396.1
AlCl	261.4
As	235.0
C <sub>2</sub>	516.5
CF <sub>2</sub>	251.9
Cl	741.4
CN	289.8, 304.2, 387.0
CO	292.5, 302.8, 313.8, 325.3, 482.5, 483.5, 519.8
F	703.7, 712.8
Ga	417.2
H	486.1, 656.5
In	325.6
N	674.0
N <sub>2</sub>	315.9, 337.1
NO	247.9, 288.5, 289.3, 303.5, 304.3, 319.8, 320.7, 337.7, 338.6
O	777.2, 844.7
OH	281.1, 306.4, 308.9
S	469.5
Si	288.2
SiCl	287.1
SiF	440.1, 777.0

<sup>1</sup> From Herman, Irving, P. (1996), "Optics Diagnostics for Thin Film Processing," Academic Press, p. 177. Ref. From Singer (1988), Selwyn (1993) and other sources cited in the text.

# Common Spectral Lines in Etching Applications

200nm	201.2 – Au	240.0 – CF	254.4 – Cl <sub>2</sub>	264.8 – AlCl	271.1 – CF <sub>2</sub>	281.0 – SiCl	288.2 – Si	
	202.1 – Au	241.4 – SiO	255.1 – W	264.7 – Ta	271.5 – Ta	281.1 – OH	288.5 – NO	
	205.5 – Cr	242.8 – Au	255.3 – P	265.1 – Ge	272.2 – BCl	281.4 – N <sub>2</sub>	289.3 – NO	
	206.8 – Ge	247.4 – CF	255.8 – CF	265.9 – Pt	272.2 – NO	282.0 – N <sub>2</sub>	289.8 – CN	
	221.6 – Ni	247.9 – NO	256.1 – Cl <sub>2</sub>	266.5 – BCl	272.4 – W	282.4 – SiCl	290.0 – CO <sub>2</sub>	
	228.8 – Cd	248.3 – Fe	257.8 – He	266.9 – SiO	274.8 – Au	283.3 – CO	292.1 – CF <sub>2</sub>	
	229.9 – SiO	248.7 – SiO	258.0 – CCl	267.6 – Au	275.0 – CF <sub>2</sub>	283.3 – Pb	292.5 – CO	
	232.0 – Ni	248.8 – CF <sub>2</sub>	259.5 – CF <sub>2</sub>	268.0 – NO	277.8 – CCl	283.7 – C	293.0 – Pt	
	234.4 – SiO	249.8 – B	259.6 – NO	268.3 – AlCl	278.8 – CCl	286.0 – BCl	295.3 – N <sub>2</sub>	
	235.0 – As	251.9 – CF <sub>2</sub>	261.4 – AlCl	269.4 – SiO	280.0 – CF <sub>2</sub>	286.0 – NO	296.2 – N <sub>2</sub>	
	237.0 – NO	253.6 – P	262.8 – Pt	269.8 – CO	280.2 – Pb	287.1 – SiCl	297.7 – N <sub>2</sub>	
	238.9 – CO	253.7 – Hg	262.9 – CF <sub>2</sub>	270.2 – Pt	280.7 – SiCl	288.0 – CO <sub>2</sub>		
	300nm	301.2 – Ni	311.7 – N <sub>2</sub>	327.4 – Cu	338.3 – Ag	358.1 – Fe	371.1 – N <sub>2</sub>	389.0 – Zr
		301.2 – Ta	312.3 – Au	328.1 – Ag	338.6 – NO	358.4 – NO	372.0 – Fe	390.2 – SiCl
302.1 – OH		313.2 – Mo	328.5 – N <sub>2</sub>	339.2 – Zr	358.6 – CN	373.7 – Fe	390.3 – Mo	
302.8 – CO		313.4 – CO	330.6 – CO	340.5 – Pd	359.0 – CN	375.5 – N <sub>2</sub>	391.2 – O	
303.5 – NO		313.6 – N <sub>2</sub>	330.9 – N <sub>2</sub>	341.5 – Ni	359.3 – Cr	376.8 – Gd	394.3 – N <sub>2</sub>	
304.2 – CN		313.8 – CO	331.1 – Ta	343.8 – Zr	360.1 – Zr	379.8 – Mo	394.4 – Al	
304.3 – NO		315.9 – N <sub>2</sub>	333.9 – N <sub>2</sub>	346.2 – Ni	360.5 – Cr	380.5 – N <sub>2</sub>	395.5 – SiF <sub>2</sub>	
306.4 – OH		317.0 – Mo	334.6 – SiF	346.6 – Cd	361.0 – Pd	382.0 – He	396.0 – Al	
306.5 – Pt		319.3 – Mo	335.0 – Gd	349.3 – CO	361.1 – Cd	385.8 – N <sub>2</sub>	396.5 – He	
306.7 – OH		319.8 – NO	335.0 – N <sub>2</sub>	349.6 – Zr	363.5 – Pd	386.2 – CN	397.3 – O	
307.0 – CCl		320.7 – NO	335.0 – Ti	350.0 – N <sub>2</sub>	363.5 – Ti	386.4 – Mo	399.8 – N <sub>2</sub>	
307.8 – OH		321.4 – CF <sub>2</sub>	336.0 – NH	352.5 – Ni	364.2 – N <sub>2</sub>	387.0 – CN	399.9 – Ti	
308.2 – Al		324.7 – Cu	336.3 – SiF	353.7 – N <sub>2</sub>	365.0 – Hg	387.1 – CN		
308.9 – OH		325.3 – CO	337.0 – CO <sub>2</sub>	357.2 – NO	365.3 – Ti	388.3 – CN		
309.3 – Al		325.6 – In	337.1 – N <sub>2</sub>	357.7 – N <sub>2</sub>	367.2 – N <sub>2</sub>	388.9 – He		
310.4 – N <sub>2</sub>		326.8 – N <sub>2</sub>	337.7 – NO	357.9 – Cr	368.3 – Pb	389.5 – N <sub>2</sub>		
400nm		400.0 – Ti	407.7 – TiF	418.3 – TiCl	425.9 – AlH	436.8 – SiF	464.9 – O	483.5 – CO
		400.9 – W	408.1 – Zr	419.0 – O	426.7 – C	438.8 – He	468.8 – Zr	486.1 – H
		402.6 – He	408.7 – SiN	419.3 – TiCl	427.0 – N <sub>2</sub>	440.1 – SiF	469.0 – Zr	488.0 – Ar
	404.7 – Hg	409.5 – N <sub>2</sub>	419.7 – CN	427.5 – Cr	440.7 – SiN	469.5 – S	489.7 – Cl	
	405.1 – SiN	410.2 – In	420.0 – N <sub>2</sub>	429.5 – W	444.3 – SiN	471.3 – He	492.2 – He	
	405.8 – Pb	411.6 – SiN	420.4 – SiN	431.4 – CH	447.1 – He	473.7 – C <sub>2</sub>		
	405.9 – N <sub>2</sub>	412.7 – SiN	421.6 – CN	434.0 – H	451.1 – In	474.2 – Ge		
	407.2 – W	414.2 – N <sub>2</sub>	423.9 – SiN	434.5 – N <sub>2</sub>	451.1 – CO	476.5 – Ar		
	407.4 – W	417.2 – Ga	424.1 – AlH	434.8 – Ar	460.0 – CO	479.4 – Cl		
	407.6 – O	418.1 – CN	425.4 – Cr	435.8 – Hg	460.2 – P	482.5 – CO		
	500nm	501.6 – He	508.6 – Cd	520.9 – Ag	546.0 – Hg	585.8 – Cn	590.6 – N <sub>2</sub>	
504.1 – Si		516.5 – C <sub>2</sub>	521.8 – Cl	561.0 – CO	585.4 – N <sub>2</sub>	595.9 – N <sub>2</sub>		
504.8 – He		519.8 – CO	542.3 – Cl	575.5 – N <sub>2</sub>	587.6 – He			
505.5 – Si		520.8 – Cr	546.6 – Ag	580.4 – N <sub>2</sub>	589.3 – Ge			
600nm		601.4 – N <sub>2</sub>	623.9 – F	639.7 – Ga	646.9 – N <sub>2</sub>	662.4 – N <sub>2</sub>	683.4 – F	696.5 – Ar
	607.0 – N <sub>2</sub>	632.3 – N <sub>2</sub>	641.3 – Ga	647.8 – CN	667.8 – He	685.4 – F	696.6 – F	
	608.0 – CO	634.7 – Si	641.4 – F	654.5 – N <sub>2</sub>	670.5 – N <sub>2</sub>	685.6 – F		
	615.6 – O	634.8 – F	643.8 – Cd	656.3 – H	674.0 – N	687.0 – F		
	615.7 – O	637.1 – Si	645.6 – O	656.5 – H	677.4 – F	690.2 – F		
	615.8 – O	639.5 – N <sub>2</sub>	646.0 – P	662.0 – CO	678.9 – N <sub>2</sub>	691.0 – F		
	700nm	703.7 – F	720.2 – F	732.6 – C	741.4 – Cl	751.5 – F	763.5 – Ar	777.2 – O
706.5 – He		725.4 – O	733.2 – F	742.6 – F	755.2 – F	772.4 – Ar	780.0 – F	
706.7 – Ar		725.6 – Cl	738.4 – Ar	750.4 – Ar	757.3 – F	775.3 – N <sub>2</sub>	787.3 – CN	
712.8 – F		727.3 – N <sub>2</sub>	738.7 – N <sub>2</sub>	750.4 – N <sub>2</sub>	760.7 – F	775.5 – F	789.6 – N <sub>2</sub>	
716.5 – N <sub>2</sub>		728.1 – He	739.9 – F	751.5 – Ar	762.6 – N <sub>2</sub>	777.0 – SiF	794.8 – Ar	
800nm	844.7 – O							

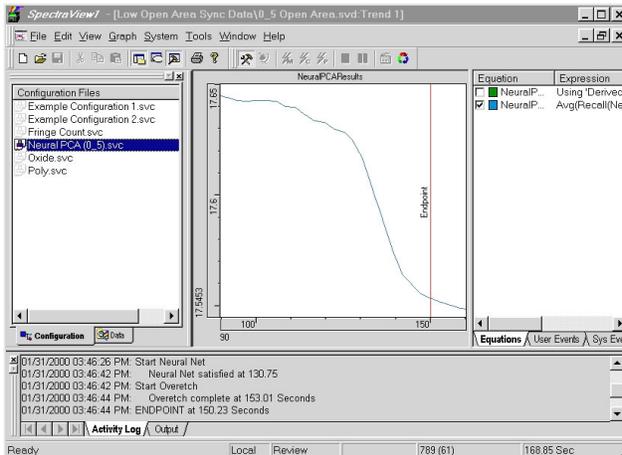
<sup>1</sup>This table is based on Verity research; Herman, Irving P. (1996), "Optics Diagnostics for Thin Film Processing," Academic Press, p. 177 (from Singer (1988), Selwyn (1993), and other sources cited in the text "CRC Handbook of Chemistry and Physics," CRC Press, Inc. (1978); and Pearse, R.W.B. and Gaydon, A. G. (1976), "Identification of Molecular Spectra", Chapman and Hall.

# Common Spectral Lines in Etching Applications

<b>A</b>	Ag	328.1, 338.3, 520.9, 546.6
	Al	308.2, 309.3, 394.4, 396.1
	AlCl	261.4, 264.8, 268.3
	AlH	424.1, 425.9
	Ar	434.8, 476.5, 488.0, 696.5, 706.7, 738.4, 750.4, 751.5, 763.5, 772.4, 794.8
	As	235.0
	Au	201.2, 202.1, 242.8, 267.6, 274.8, 312.3
<b>B</b>	B	249.8
	BCl	266.5, 272.2, 286.0
<b>C</b>	C	283.7, 426.7, 732.6
	C <sub>2</sub>	473.7, 516.5
	CCl	258.0, 277.8, 278.8, 307.0, 460.0
	Cd	228.8, 346.6, 361.1, 508.6, 643.8
	CF	240.0, 247.4, 255.8
	CF <sub>2</sub>	248.8, 251.9, 259.5, 262.9, 271.1, 275.0, 280.0, 292.1, 321.4
	CH	431.4
	Cl	479.4, 489.7, 521.8, 542.3, 725.6, 741.4
	Cl <sub>2</sub>	254.4, 256.1
	CN	289.8, 304.2, 358.6, 359.0, 386.2, 387.0, 387.1, 388.3, 418.1, 419.7, 421.6, 585.8, 647.8, 787.3
	CO	238.9, 269.8, 283.3, 292.5, 302.8, 313.4, 313.8, 325.3, 330.6, 349.3, 451.1, 482.5, 483.5, 519.8, 561.0, 608.0, 662.0
	CO <sub>2</sub> <sup>+</sup>	337.0
	Cr	205.5, 357.9, 359.3, 360.5, 425.4, 427.5, 520.8
Cu	324.7, 327.4	
<b>F</b>	F	623.9, 634.8, 641.4, 677.4, 683.4, 685.4, 685.6, 687.0, 690.2, 691.0, 696.6, 703.7, 712.8, 720.2, 733.2, 739.9, 742.6, 755.2, 757.3, 760.7, 775.5, 780.0
	Fe	248.3, 358.1, 372.0, 373.7
<b>G</b>	Ga	417.2, 639.7, 641.3
	Gd	335.0, 376.8
	Ge	206.8, 265.1, 474.2, 589.3
<b>H</b>	H	434.0, 486.1, 656.3, 656.5
	He	257.8, 382.0, 388.9, 396.5, 402.6, 438.8, 447.1, 471.3, 492.2, 501.6, 504.8, 587.6, 667.8, 706.5, 728.1
	Hg	253.7, 365.0, 404.7, 435.8, 546.1
<b>I</b>	In	325.6, 410.2, 451.1
<b>M</b>	Mo	313.2, 317.0, 319.3, 379.8, 386.4, 390.3
<b>N</b>	N	674.0
	N <sub>2</sub>	281.4, 282.0, 295.3, 296.2, 297.7, 310.4, 311.7, 313.6, 315.9, 326.8, 328.5, 330.9, 333.9, 335.0, 337.1, 350.0, 353.7, 357.7, 364.2, 367.2, 371.1, 375.5, 380.5, 385.8, 389.5, 394.3, 399.8, 405.9, 409.5, 414.2, 420.0, 427.0, 434.5, 575.5, 580.4, 585.4, 590.6, 595.9, 601.4, 607.0, 632.3, 639.5, 646.9, 654.5, 662.4, 670.5, 678.9, 716.5, 727.3, 738.7, 750.4, 762.6, 775.3, 789.6
	NH	336.0
	Ni	221.6, 232.0, 301.2, 341.5, 346.2, 352.5
	NO	237.0, 247.9, 259.6, 268.0, 272.2, 286.0, 288.5, 289.3, 303.5, 304.3, 319.8, 320.7, 337.7, 338.6, 357.2, 358.4
	O	391.2, 397.3, 407.6, 419.0, 464.9, 615.6, 615.7, 615.8, 645.6, 725.4, 777.2, 844.7
OH	281.1, 302.1, 306.4, 306.7, 307.8, 308.9	
<b>P</b>	P	253.6, 255.3, 460.2, 646.0
	Pb	280.2, 283.3, 368.3, 405.8
	Pd	340.5, 361.0, 363.5
	Pt	262.8, 265.9, 270.2, 293.0, 306.5
<b>S</b>	S	469.5
	Si	288.2, 504.1, 505.5, 634.7, 637.1
	SiCl	280.7, 281.0, 282.4, 287.1
	SiF	334.6, 336.3, 436.8, 440.1, 777.0
	SiF <sub>2</sub>	390.2, 395.5
	SiN	405.1, 408.7, 411.6, 412.7, 420.4, 423.9, 440.7, 444.3
SiO	229.9, 234.4, 241.4, 248.7, 266.9, 269.4	
<b>T</b>	Ta	264.7, 271.5, 301.2, 331.1
	Ti	335.0, 363.5, 365.3, 399.9
	TiCl	418.3, 419.3
	TiF	407.7
<b>W</b>	W	255.1, 272.4, 400.9, 407.4, 429.5
<b>Z</b>	Zn	339.2, 343.8, 349.6, 360.1, 389.0, 407.3, 408.1, 468.8

# Algorithms for Endpoint Detection

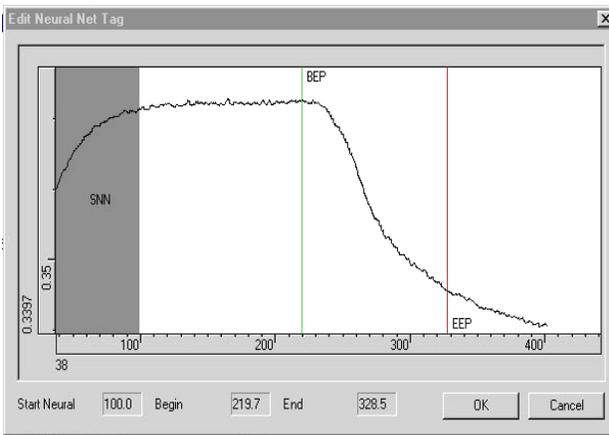
Verity Instruments provides a powerful suite of endpoint algorithms, including the multivariate Neural PCA algorithm, which can be processed with Verity's proprietary Neural Network pattern recognition software.



## Neural PCA Multi-Wavelength Algorithm

Verity's endpoint-detection computations can employ robust algorithms such as the patent pending Neural PCA for multivariate, full-spectrum analysis.

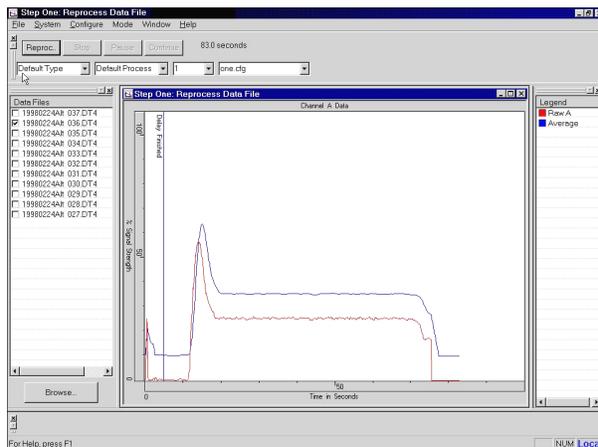
Data represented at left apply to a 0.5% exposed area contact etch using Verity's Neural PCA algorithm. Within the SpectraView™ endpoint software application, the Neural PCA endpoint trace can be processed using Neural Network or threshold-based methods.



## Neural Network Algorithm

The Neural Network algorithm is used to analyze endpoint traces. The Neural Network uses proprietary techniques to recognize characteristic endpoint shapes in the trend line. This is performed in real time and the pattern recognition algorithm adapts to expected amplitude and duration changes in the endpoint trace during successive runs.

Unlike other types of neural networks, Verity's algorithm can be set up with only a few training runs. If a false positive or negative is found, it can easily be added to the training set for improved robustness. Process engineers using Verity's Neural Net software are freed from the burden associated with developing and testing threshold-based algorithms. In addition, data can be analyzed "on-the-fly" or replayed, reviewed, or reprocessed with SpectraView™.



## Threshold-Based Algorithm

Using threshold-based algorithms, endpoint recognition is based upon the output rising above or below a preset level for a predetermined length of time. However, for demanding applications, the Neural Network algorithm is commonly selected over the threshold-based algorithm.

## Suggested Further Reading

Applied Optics and Optical Engineering, Volume V, Optical Instruments, Part II. 1969. Academic Press, New York and London. (For Design, Construction and Testing of Optical Devices)

"CRC Handbook of Chemistry and Physics – A Ready-Reference Book of Chemical and Physical Data." 1978. *CRC Press*. (Scientific Reference of Atomic Spectra)

Herman, Irving P. 1996. "Optic Diagnostics For Thin Film Processing", *Academic Press*, ch. 6. (Optical Emission Spectroscopy as Applied to Thin Film Processing)

Pearse, R.W.B. and Gaydon, A.G. 1976. "The Identification of Molecular Spectra", *Chapman and Hall*. (Scientific Reference of Molecular Spectra)

Selwyn, Gary S. 1993. "Optical Techniques for Plasma Processing", *American Vacuum Society*. (A "How to" Book to Enable Better Understanding of Optical Techniques)